## Length of meridian arc on an ellipsoid

Suppose the Earth is modelled as an ellipsoid with semi-major axis a, semi-minor axis b, and eccentricity e. A standard problem is to find the length of a meridian arc between two geographical latitudes, say  $\varphi_1$  and  $\varphi_2$ . Three methods are suggested below. For more on this topic see e.g. the lecture notes by Deakin and Hunter [1].

The radius of curvature in meridian at latitude  $\varphi$  is given by

$$\rho = a(1 - e^2)(1 - e^2\sin^2\varphi)^{-3/2} = a(1 + \varepsilon)^{1/2}(1 + \varepsilon\cos^2\varphi)^{-3/2}$$

where 
$$\varepsilon = e^2/(1 - e^2) = (a^2 - b^2)/b^2$$
.

We work here in terms of  $\varepsilon$ , which is assumed to be small ( $\varepsilon \approx 1/150$  for the Earth). The arc length is given by

$$a(1+\varepsilon)^{1/2}\int_{\varphi_1}^{\varphi_2}(1+\varepsilon\cos^2\varphi)^{-3/2}d\varphi.$$

# Method 1: Numerical integration

In theory this method is subject to rounding errors, due to the addition of many terms. In practice it seems to work well enough, provided one uses an accurate integration formula such as Weddle's rule (rather than the better known Simpson's and three-eighths rules). In Weddle's rule the number of intervals is a multiple of 6, and the integral over a block of 6 intervals is estimated as

$$\int_0^{6h} y(x) dx = \frac{3h}{10} [y(0) + 5y(h) + y(2h) + 6y(3h) + y(4h) + 5y(5h) + y(6h)].$$

The number of intervals can be repeatedly doubled until the value stays constant within the desired order of accuracy. This method is slow, but is useful as a check on other methods.

### Method 2: integration of the power series

Here we expand the integrand  $(1 + \varepsilon \cos^2 \varphi)^{-3/2}$  by the binomial theorem, taking as many terms as are required to achieve the desired accuracy. The powers  $\cos^{2m} \varphi$  are integrated by rewriting them in terms of multiple angles:

$$\cos^{2m}\varphi = 2^{1-2m} \left[ \frac{1}{2} {2m \choose m} + {2m \choose m-1} \cos 2\varphi + {2m \choose m-2} \cos 4\varphi + {2m \choose m-3} \cos 6\varphi + \dots + \cos 2m\varphi \right]$$

We may as well multiply the result by the binomial expansion of  $(1 + \varepsilon)^{1/2}$ , since this does not make it any more complicated. If this is done, the length of the meridian arc is

$$a[c_0(\varphi_2 - \varphi_1) + c_1(\sin 2\varphi_2 - \sin 2\varphi_1) + c_2(\sin 4\varphi_2 - \sin 4\varphi_1) + c_3(\sin 6\varphi_2 - \sin 6\varphi_1) + \dots],$$

where the  $c_m$  are power series in  $\varepsilon$ . To reduce the effect of rounding errors when  $\varphi_2 - \varphi_1$  is small, it may be better to rewrite  $\sin 2m\varphi_2 - \sin 2m\varphi_1$  as  $2\sin m(\varphi_2 - \varphi_1)\cos m(\varphi_2 + \varphi_1)$ . The coefficients  $c_m$  are given below up to  $\varepsilon^8$  (equivalent to  $e^{16}$ ), which should be more than enough for most purposes.

$$c_0 = 1 - \frac{1}{4}\varepsilon + \frac{13}{64}\varepsilon^2 - \frac{45}{256}\varepsilon^3 + \frac{2577}{16384}\varepsilon^4 - \frac{9417}{65536}\varepsilon^5 + \frac{139613}{1048576}\varepsilon^6 - \frac{522821}{4194304}\varepsilon^7 + \frac{126287705}{1073741824}\varepsilon^8 - \dots$$

$$c_1 = -\frac{3}{8}\varepsilon + \frac{9}{32}\varepsilon^2 - \frac{237}{1024}\varepsilon^3 + \frac{819}{4096}\varepsilon^4 - \frac{23325}{131072}\varepsilon^5 + \frac{84711}{524288}\varepsilon^6 - \frac{4993233}{33554432}\varepsilon^7 + \frac{18593103}{134217728}\varepsilon^8 - \dots$$

$$c_2 = \frac{15}{256}\varepsilon^2 - \frac{75}{1024}\varepsilon^3 + \frac{1245}{16384}\varepsilon^4 - \frac{4905}{65536}\varepsilon^5 + \frac{607125}{8388608}\varepsilon^6 - \frac{2332785}{33554432}\varepsilon^7 + \frac{35785995}{536870912}\varepsilon^8 - \dots$$

$$c_3 = -\frac{35}{3072}\varepsilon^3 + \frac{245}{12288}\varepsilon^4 - \frac{19985}{786432}\varepsilon^5 + \frac{90475}{3145728}\varepsilon^6 - \frac{3093755}{100663296}\varepsilon^7 + \frac{12812765}{402653184}\varepsilon^8 - \dots$$

$$c_4 = \frac{315}{131072}\varepsilon^4 - \frac{2835}{524288}\varepsilon^5 + \frac{68859}{8388608}\varepsilon^6 - \frac{354627}{33554432}\varepsilon^7 + \frac{26776197}{2147483648}\varepsilon^8 - \dots$$

$$c_5 = -\frac{693}{1310720}\varepsilon^5 + \frac{7623}{5242880}\varepsilon^6 - \frac{430353}{167772160}\varepsilon^7 + \frac{498267}{134217728}\varepsilon^8 - \dots$$

$$c_6 = \frac{1001}{8388608}\varepsilon^6 - \frac{13013}{33554432}\varepsilon^7 + \frac{418847}{536870912}\varepsilon^8 - \dots$$

$$c_7 = -\frac{6435}{234881024}\varepsilon^7 + \frac{96525}{939524096}\varepsilon^8 - \dots$$

$$c_8 = \frac{109\,395}{17\,179\,869\,184} \varepsilon^8 - \dots$$

## Method 3: Another power series method

The following method may be of interest in that it can be extended by computer to any desired degree of accuracy, without the need to program-in complicated coefficients like those in Method 2. It finds the arc length along the meridian from the equator to a given latitude  $\varphi$ . The arc length from  $\varphi_1$  to  $\varphi_2$  can be found by applying the method to  $\varphi_1$  and  $\varphi_2$ , and taking the difference.

We expand  $(1 + \epsilon \cos^2 \varphi)^{-3/2}$  by the binomial theorem, as in Method 2, but now integrate the  $\cos^{2m} \varphi$  as follows. Setting

$$K_m = \int_0^{\varphi} \cos^{2m} \varphi d\varphi,$$

we find easily that  $2mK_m = (2m-1)K_{m-1} + \sin\varphi \cos^{2m-1}\varphi$ , whence

$$K_0 = \varphi$$
,  $K_1 = \frac{1}{2}\varphi + \frac{1}{2}\sin\varphi\cos\varphi$ ,  $K_2 = \frac{3}{8}\varphi + \frac{3}{8}\sin\varphi\cos\varphi + \frac{1}{4}\sin\varphi\cos^3\varphi$ , etc.

A routine calculation shows that the integral, which has to be multiplied by  $a(1 + \varepsilon)^{1/2}$  (=  $a^2/b$ ) to get the arc length, is

$$\phi \left(1 - \frac{3}{4}\epsilon + \frac{3}{4} \cdot \frac{15}{16}\epsilon^{2} - \frac{3}{4} \cdot \frac{15}{16} \cdot \frac{35}{36}\epsilon^{3} + \frac{3}{4} \cdot \frac{15}{16} \cdot \frac{35}{36} \cdot \frac{63}{64}\epsilon^{4} - \frac{3}{4} \cdot \frac{15}{16} \cdot \frac{35}{36} \cdot \frac{63}{64} \cdot \frac{99}{100}\epsilon^{5} + \dots \right)$$

$$+ \sin\phi\cos\phi \left( -\frac{3}{4}\epsilon + \frac{3}{4} \cdot \frac{15}{16}\epsilon^{2} - \frac{3}{4} \cdot \frac{15}{16} \cdot \frac{35}{36}\epsilon^{3} + \frac{3}{4} \cdot \frac{15}{16} \cdot \frac{35}{36} \cdot \frac{63}{64}\epsilon^{4} - \frac{3}{4} \cdot \frac{15}{16} \cdot \frac{35}{36} \cdot \frac{63}{64} \cdot \frac{99}{100}\epsilon^{5} + \dots \right)$$

$$+ \frac{1}{2}\sin\phi\cos^{3}\phi \left( \frac{15}{16}\epsilon^{2} - \frac{15}{16} \cdot \frac{35}{36}\epsilon^{3} + \frac{15}{16} \cdot \frac{35}{36} \cdot \frac{63}{64}\epsilon^{4} - \frac{15}{16} \cdot \frac{35}{36} \cdot \frac{63}{64} \cdot \frac{99}{100}\epsilon^{5} + \dots \right)$$

$$+ \frac{1}{2} \cdot \frac{3}{4}\sin\phi\cos^{5}\phi \left( \frac{3}{2}\epsilon^{3} + \frac{3}{2}\epsilon^{3} \cdot \frac{63}{36}\epsilon^{3} + \frac{3}{2}\epsilon^{3} \cdot \frac{63}{64}\epsilon^{4} - \frac{3}{2}\epsilon^{3} \cdot \frac{63}{64}\epsilon^{4} - \frac{3}{2}\epsilon^{3} \cdot \frac{63}{64}\epsilon^{4} - \frac{3}{2}\epsilon^{3} \cdot \frac{63}{64}\epsilon^{4} - \frac{99}{100}\epsilon^{5} + \dots \right)$$

$$+ \frac{1}{2} \cdot \frac{3}{4} \cdot \frac{5}{6}\sin\phi\cos^{7}\phi \left( \frac{63}{2}\epsilon^{3} + \frac{3}{2}\epsilon^{3} \cdot \frac{63}{64}\epsilon^{4} - \frac{3}{2}\epsilon^{3} \cdot \frac{63}{64}\epsilon^{4} - \frac{3}{2}\epsilon^{3} \cdot \frac{63}{64}\epsilon^{4} - \frac{3}{2}\epsilon^{3} \cdot \frac{63}{64}\epsilon^{4} - \frac{99}{100}\epsilon^{5} + \dots \right)$$

$$+ \frac{1}{2} \cdot \frac{3}{2} \cdot \frac{5}{4} \cdot \frac{5}{6}\sin\phi\cos^{7}\phi \left( \frac{63}{2}\epsilon^{3} + \frac{3}{2}\epsilon^{3} \cdot \frac{63}{64}\epsilon^{4} - \frac{3}{2}\epsilon^{3} \cdot \frac{63}{64}\epsilon^{4} - \frac{3}{2}\epsilon^{3} \cdot \frac{63}{64}\epsilon^{4} - \frac{99}{100}\epsilon^{5} + \dots \right)$$

$$+ \frac{1}{2} \cdot \frac{3}{2} \cdot \frac{5}{4} \cdot \frac{7}{6}\sin\phi\cos^{7}\phi \left( \frac{63}{2}\epsilon^{3} + \frac{3}{2}\epsilon^{3} \cdot \frac{63}{64}\epsilon^{4} - \frac{3}{2}\epsilon^{3} \cdot \frac{63}{64}\epsilon^{4} - \frac{3}{2}\epsilon^{3} \cdot \frac{63}{64}\epsilon^{4} - \frac{99}{100}\epsilon^{5} + \dots \right)$$

$$+ \frac{1}{2} \cdot \frac{3}{2} \cdot \frac{5}{4} \cdot \frac{7}{6}\sin\phi\cos^{9}\phi \left( \frac{3}{2}\epsilon^{3} + \frac{3}{2}\epsilon^{3} \cdot \frac{15}{6}\epsilon^{3} + \frac{3}{2}\epsilon^{3} \cdot \frac{15}{6}\epsilon^{3} + \frac{3}{2}\epsilon^{3} + \frac{3$$

It is clear how the pattern can be continued indefinitely. The following tested piece of code, written in Delphi Pascal, shows how the method can be applied to give any degree of precision up to the limit of the language used.

```
// arbitrary: higher if machine precision warrants it
  coeff : array [0..MAX_POWER_EPSILON] of extended;
  // Delphi "extended" is floating-point with 64-bit precision (19 sig. fig.)
Compute the coefficients of phi, sin(phi)*cos(phi), sin(phi)*cos^3(phi), etc.
This procedure need only be called once for a given ellipsoid (a, b).
procedure ComputeCoefficients( a, b : extended);
  epsilon : extended;
  j, k : integer;
 denom, mult : extended;
  epsilon := (a - b)*(a + b) / (b*b);
  // Compute sums shown inside large () in Web text
  j := MAX_POWER_EPSILON;
  denom := 4.0*j*j;
  coeff[j] := (denom - 1.0)/denom;
  for j := MAX_POWER_EPSILON - 1 downto 0 do begin
    for k := MAX_POWER_EPSILON downto j + 1 do coeff[k] := -epsilon * coeff[k];
    if (j > 0) then begin
      denom := 4.0*j*j;
      coeff[j] := (coeff[j+1] + 1.0) * (denom - 1.0)/denom;
    else coeff[0] := coeff[1] + 1.0;
  end;
```

```
// Apply factors 1/2, 1.3/2.4, 1.3.5/2.4.6, etc. and dimensions of ellipsoid.
  mult := a*a/b;
  coeff[0] := mult*coeff[0];
  coeff[1] := mult*coeff[1];
  for j := 2 to MAX_POWER_EPSILON do begin
    mult := mult*(2*j - 3)/(2*j - 2);
    coeff[j] := mult*coeff[j];
 end;
end;
Apply the above coefficients to compute arc length on meridian
 from equator to passed-in latitude phi.
function ComputeArcLengthFromEquator( phi_degrees : extended) : extended;
 phi, c, c2, total : extended;
 j : integer;
begin
 phi := (PI/180.0) * phi_degrees; // degrees to radians
  c := Cos( phi);
 c2 := c*c;
  total := coeff[MAX_POWER_EPSILON];
 for j := MAX_POWER_EPSILON - 1 downto 1 do total := total*c2 + coeff[j];
 result := phi*coeff[0] + Sin(phi)*c*total;
```

#### Reference

[1] R.E. Deakin and M.N. Hunter, *Geometric Geodesy A*. Downloaded from <a href="http://user.gs.rmit.edu.au/rod/files/publications/Geometric%20Geodesy%20A(2010).pdf">http://user.gs.rmit.edu.au/rod/files/publications/Geometric%20Geodesy%20A(2010).pdf</a>